# IONOSPHERIC SCINTILLATIONS INDUCED BY CLOUD OF INTENSE SPORADIC E LAYER

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## ABSTRACT

The simultaneous observations of vertical ionospheric soundings and the faraday fadings of the beacons from a low orbiting sitellite have shown clearly that the existence of a cloud of intense sporadic E in the path of radio waves does produce scintillations even during the day time at low latitudes. The latitudinal width of E<sub>s</sub> cloud is estimated to be around 300 km.

# Introduction

EARLY studies of ionospheric scintillations were carried out at middle and high latitudes using the signals from radio stars which traversed the whole ionosphere. The signals received were found to fade violently during the night time. This phenomenon known as ionospheric scintillation of radio waves was

found to be associated with the occurrence of spread F in the path of radio waves (Wild and Roberts, 1956; Booker, 1958).

With the advent of artificial earth satellites, scintillation studies have been extended to all latitudes. In middle latitudes again, a positive correlation was found between scintillations and spread F (Yeh and

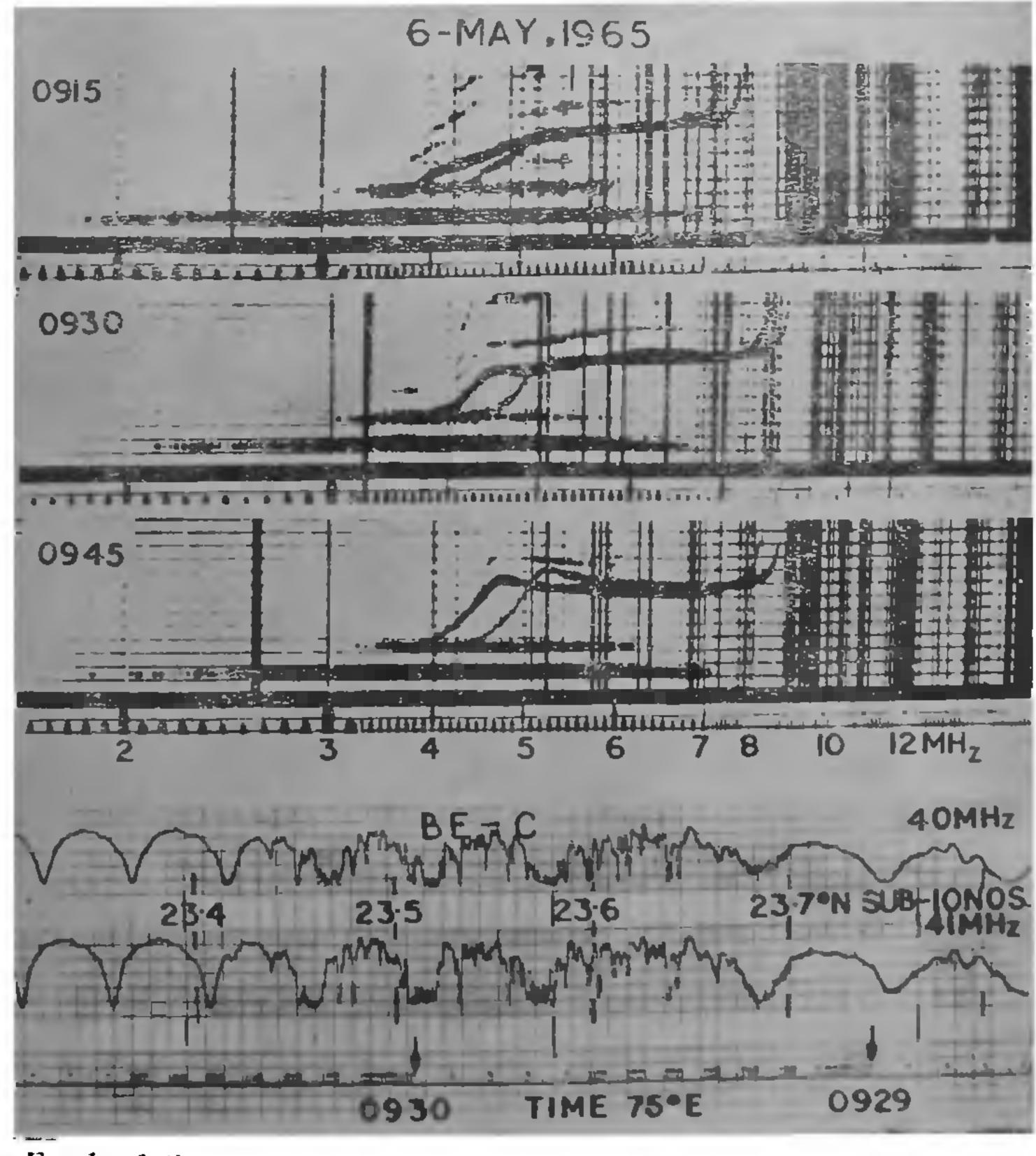


Fig. 1. Faraday fading records at 40 and 41 MHz from the Bl. C Satellite recorded at Ahmedabad on 6 May 1965 showing scintillations (bottom), caused by the intense sporadic E scen in the ionograms at the top.

Swenson, 1959; Beynon and Jones, 1964; McClure, 1964). At equatorial latitudes too, the scintillations were found to be a night time phenomenon associated with spread F (Kent, 1961; Koster, 1972; Chandra and Rastogi, 1974).

(1968) showed that the day time scintillations observed at Sagamore Hill (dip 73° N) were correlated with the occurrence of E<sub>s</sub>, the width of E<sub>s</sub> cloud estimated by them was 300-600 km. At a low latitude station, Brisbane (20° N), Singleton and Lynch (1962) did not find any significant correlation between scintillations and foE<sub>8</sub>. Koster and Wright (1963) have observed an unusually regular type of scintillation during day time at Accra (dip 8-5°S) which was associated with equatorial electrojet disturbances, Ramakrishnan (1966) has reported the associations between the occurrences of scintillations and E<sub>s</sub>. The present note is concerned with a conclusive evidence for stroeng blanketing type of E<sub>s</sub> causing scintillations at 40 and 41 MHz.

### RESULTS

In Fig. 1 are shown the ionograms taken at Ahmedabad at 0915, 0930 and 0945 hr 75° EMT on 6 May 1965. All these records show a strong E<sub>s</sub> cloud over Ahmedabad with fE<sub>s</sub> > 7.0 MHz, and the second multiple of E<sub>s</sub> is also clearly seen to be strong. In the lower part of the diagram are shown the fading records of 40 and 41 MHz radio beacons from the Satellite BE-C recorded at Ahmedabad. The subsatellite pass was very close to the location of Ahmedabad, the maximum elevation of the satellite was 82° at 0930 hr 75° EMT. The fading records for periods prior to 0929 hr, i.e., for sub-ionospheric position north

of 23.7° N showed very clear Faraday fades. Between 09h, 29', 12" and 09h, 30' 20" strong scintillations were superimposed over the Faraday fades. The sub-ionospheric latitude varied from 23.7° N to 23.4° N during this period. After 0931 hr, when the satellite was At high latitude stations scintillations have been south of 23.4° N the Faraday fades were again quite found to oecur some times during the day and are clear. This showed that strong irregularities were attributed to E-region irregularities (Bolton et al., present in the ionospheric between latitues of 23.4° N 1953; Dueno, 1956; Chivers and Greenhow, 1959; and 23.7°N. This clearly shows a distinct correlation Munro, 1966). Using spaced receivers McClure (1964) of strong blanketing E, and the scintillations produced estimated the height of irregularities causing the day in the radio waves traversing through it. The E, time scintillations to be at 100 km. Aarons and Whitney cloud seems to be about 300 Km wide in N-S direction.

#### ACKNOWLEDGEMENT

The research at the Physical Research Laboratory is financially supported by the Department of Space, Government of India.

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# PREPARATION AND CHARACTERISATION OF ORTHOPHENYLENEDIAMINE COMPLEXES OF LANTHANIDES

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RTHOPHENYLENEDIAMINE (OPDA) com plexes of Fe (II) halides were prepared and characterised by infrared, electronic, Mössbauer

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spectroscopy and magnetic susceptibility measure. ments<sup>1</sup>. Some complexes of Co(II), Ni(II) and Cu (II) with orthophenylenediamine have been prepared and characterized by magnetic and spectral properties2. In all the above complexes it was observed OPDA acts either as a bidentate ligand or as a unidentate ligand.

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